School of Mathematics, Meteorology

Symmetry-break mixing, instability, and low frequency variability i[n a minimal](#page-1-0) Lorenz like system

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properties (ergodicity, hyperbolicity) and that cause the model losing ability in describing intrinsically multiscale processes.

1. Introduction

The Lorenz system [1] has a central role in modern science as it has provided the first example of low-dime

Palmer [17] introduced artificially

where g is the gravity acceleration, \vec{v} u, w $\frac{1}{z}$, $\frac{x}{x}$, T T_0 $\frac{x}{z}$ T/H , with H uniform depth of the fluid and T imposed temperature difference. The suitable boundary \tilde{w} \tilde{w} \tilde{w} See [3] for a detailed derivation in the case Ec 0, considering that the sign of the term proportional to *R* is wrong in both Eqs. 34 and 35. In the case *Ec* 0 , the seminal Lorenz system can be derived by severely truncating the system, considering the evolution equation for the real part of $\epsilon_{1,1}$ and for the imaginary parts of $\epsilon_{1,1}$ and $\epsilon_{0,2}$, and performing suitable rescaling (see below). While $_{0,2}$ has no real part because of the boundary conditions [3], neglecting the imaginary (real) part of $\mathcal{L}_{1,1}$ ($\mathcal{L}_{1,1}$) amounts to an arbitrary selection of the phase of the waves in the system. An entire hierarchy of generalized Lorenz models, all obeying to this constraint, can be derived with lengthy but straightforward calculations. See, *e.g.*, [4-6] for detailed discussion of these models.

3. Symmetries of the extended Lorenz system

In this work we include the modes $\frac{1}{1,1}$, $\frac{2,2}{1,1}$, $\frac{2,2}{1,1}$, $\frac{2,2}{1,1}$ in our truncation, and retain both the real and the imaginary parts, whereas the considered horizontally symmetric modes $\epsilon_{0.2}$ and $\epsilon_{0.4}$ are, as mentioned above, imaginary. Finally, we assume, in general, a non vanishing value for E_c . If we define X_1 *iX* X_2 , X_1 *iY* Y_1 *iY* Y_2

 $\frac{1}{\sqrt{2}}$ $\begin{array}{c} I \\ I \end{array}$

being the 2X2 rotation m

properties of Z_1 do not depend on the initial conditions, and agree with those of the *z* variable of the classical Lorenz system. Moreover, the statistical properties of the quadratic quantities Y_2^2 , Y_1^2 Y_2^2 , and 2 2 X^2 X_1^2 X_2^2 , Y^2 Y_1^2 Y_2^2 2 Y^2 Y_1^2 Y_2^2 , and XY X_1Y_2 X_2Y_1 do not depend on the initial conditions (whereas those of each term in the previous sums do!), and agree (for all values of r , b !) with those of x^2 and y^2 , and xy of the classical Lorenz system, respectively. Analogously, since the do not dep determining a chaotic dynamics for the A_1, A_2, B_1, B_2, Z_2 variables and a trivial or periodic behavior for the X_1, X_2, Y_1, Y_2, Z_1 variables.

With the "classical" Lorenz parameter values r 28, 10, b 8/3, the variables X_1, X_2, Y_1, Y_2, Z_1 obviously have an erratic behavior, whereas the variables $I_1, B_1/\sqrt{\frac{2}{M}}$ $m\sqrt{\frac{2}{M}}$ $\sqrt{\frac{2}{M}}$ which describe the faster spatially varying wave components, do not feature any time variability when asymptotic dynamics is considered, as they converge to fixed values

derivation) of the model discussed above, the fact that we recover the classical results of the 3 component Lorenz system is quite reinsuring. The Kaplan-Yorke dimension [24] of the system is:

$$
d_{KY} \quad k \quad \frac{j}{\left|\begin{array}{c} 1 \\ k & 1 \end{array}\right|} \quad 4.183 \tag{6}
$$

where *k* is such that the sum of the first *k* (4, in our case) Lyapunov exponents is positive and the sum of the fi

is beyond the scopes of this paper, as we confine ourselves to studying the properties of the system when chaotic motion is realized, thus focusing on the $Ec = 0$ limit. We consider $Ec = 0.002$.

In physical terms, the coupling allows for a mixing of the phases of the thermal and streamfunction waves: after a O^{-1} time the system "realizes" that the symmetry (4-5) is broken, so that the previously described degeneracies are destroyed and ergodicity is established in the system. Note that, since also in the case of E_c 0 the convection does not preferentially act on waves of a specific phase, we have that, pairwise, the statistical properties of $\frac{1}{1+\alpha}$ and are identical and do not depend on the initial conditions. The same applies for the and i **6.**

highlights on these variables, in Fig. 2 we show, from top to bottom, the projection on $A₁$ of three typical trajectories for $Ec \ 10^{-3}$, $Ec \ 10^{-4}$, and $Ec \ 10^{-5}$, respectively. Going from the top via the middle to the bottom panel, the time scale increases by a factor of 10 and 100, respectively. The striking geometric similarity underlines that the time scales of the dominating variability for the slow variables can be estimated as

mixing requires Ec^{-1} time units. The same considerations apply for the Y_1 and Y_2 pair of variables.

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Further insight can be obtained by looking at the spectral properties of the fast and slow variables; some relevant examples of power spectra for $Ec \ 10^{-3}$, $Ec \ 10^{-4}$, and $Ec \ 10^{-5}$ are depicted in Fig. 4, where we have considered the low frequency range by concentrating on time scales larger than $10/1$. The white noise nature of the fast variable is apparent and so is the overall lack of sensitivity of its spectral properties with respect to amplitude modulation, which is responsible for phase mixing observed in Fig. 3, is virtually absent. This shows how crucial physical processes are masked, due to their weakness, when using a specific *metric* **Exote that any signature of the** This shows how crucial properties are the specific metric of Fig. 4, where we have considered the low frequency range by concentrating on time r than $10/1$. The white noise nature of the fast variable is appearent and so is the of sensitivity of its spectral properties with respec ons apply for the Y_1 and Y_2 pair of
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 $m,n = 1,1$ and $m,n = 2,2$, and the imaginary part of the x-symmetric modes $\frac{0}{2}$ and $\frac{0}{2}$. As modes characterized by a faster varying spatial structure and different parity are added to the Lorenz spectral truncation, which describes the dynamics of the real part of $\frac{1}{11}$ and of the imaginary part of $_{1,1}$ and $_{0,2}$ only, the system we consider in this work represents the thermal impact of viscous dissipation, controlled by the Eckert number *Ec*. The presence of a forth parameter - together with the usual Lorenz parameters *r* (relative Rayleigh number), (Prandl number)*,* and *b* (geometric factor) – marks a crucial difference between this ODEs system and other extensions of the Lorenz system proposed in the literature - see, *e.g.*, [4-6]. Moreover, as can be deduced following the argumentations of Nicolis [18], this system specifically allows for the closed-form computation of the entropy production, as the Eckert number enters into its evaluation. Up to our knowledge, this is the first ODE system obtained as a truncation of the convective equations presented in the scientific literature featuring all these additional properties.

exponent coincides with that of the Lorenz system, and the value of one of the six negative exponents agrees with that of the negative Lyapunov exponent of the Lorenz system. Therefore, the Lorenz system contains already all the interesting unstable dynamics described by this extended ODEs system, and features exactly the same value for the metric entropy. Correspondingly, while the five variables (*fast*) describing the modes $_{1,1}$, $_{1,1}$, and $_{0.2}$ have an erratic behavior, the other five variables (*slow*) converge to fixed values.

2. When *Ec* 0*,* the symmetry of the system is broken, and coupling occurs between the fast and slow variable over a time scale O Ec \cdot \cdot This is clarified by adopting standard multiscale formalism. If we select, as in the original Lorenz system, $r = 28$, $= 10$, $b=8/3$, the system is chaotic for $0< Ec$ 0.045, whereas for higher values of *Ec* a quasiperiodic regime is realized. In the chaotic regime, the symmetry-break is accompanied by the establishment 12.bg9.o.7976lP AMCID ol2.0dyineShape3nv /.9450 13.0cs: two Lyapui7.76lP AMCII modulation occurring on time scales of *Ec*^{*l*} which superimposes on the fast dynamics controlled by the $O(1)$ time scale $1/$ \ldots

The system introduced in this paper features very rich dynamics and, therefore, may have prototypical value for phenomena generic to complex systems, such as the interaction between slow and fast variables and the presence of long wee2ls40 1.900.81 24 16.98 re8.15547 699.412.0108 0 0

- 2. extension of the present analysis to higher order truncation ODEs systems, with detailed investigation of the $Ec = 0$ invariance properties, of the impact on these symmetries resulting from setting $Ec > 0$, and the ensuing multiscale analysis. This is especially relevant in the context of the results presented by Franceschini and Tebaldi [26] and Franceschini et al. [27], who emphasized that spectral truncation and modes selections procedures have to be critically addressed. resulting from setting $Ez > 0$, and the ensuing multiscult unitysis. This is depending
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	- 3. definition, following [7], of the minimal truncated 3D model of convection able to represent the thermal impact of viscous dissipation;

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 (c)

Figure 1: Four largest Lyapunov exponents (a), metric entropy h (b) and Kaplan-Yorke d_{KY} dimension (c) as a **function of Ec. Note the continuity for Ec = 0 of all parameters and the distinct linear behavior - see specifically the dashed lines in (b) and (c) - for Ec < 0.008. The linear behavior of the second and fourth Lyapunov** exponents branching off zero in (a) extends throughout Ec 0.018. Details in the text.

Figure 2: Impact of the viscous-thermal feedback on the time scales of the system. From top to bottom: typical evolution of the variable A_1 for different values of the Eckert number (Ec = 10^{-3} , Ec = 10^{-4} , Ec = 10^{-5} , respectively). Note that the time scale is magnified by a factor of 1, 10 and 100 from top to bottom. Details in the **text.**

Figure 3: Impact of the viscous-thermal feedback on the time scales of the system. From top to bottom: typical evolution of the variable X_1 for different values of the Eckert number (Ec = 10⁻³, Ec = 10⁻⁴, Ec = 10⁻⁵, respectively). Note that the time scale is magnified by a factor of 1, 10 and 100 from top to bottom. Details in the **text.**

Figure 4: Power spectrum of A_1 and X_1 (multiplied times 10⁻⁸) for Ec = 10⁻³, Ec = 10⁻⁴

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